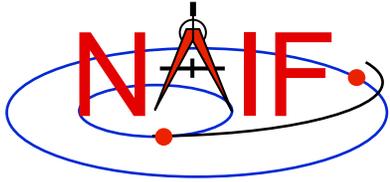


Navigation and Ancillary Information Facility

Writing a Mice (MATLAB) Based Program

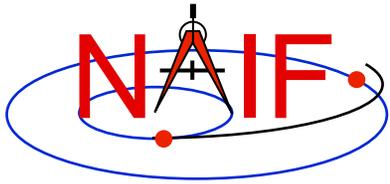
April 2016



Viewing This Tutorial

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Undefined variables are displayed in **red**
Results are displayed in **blue**



Introduction

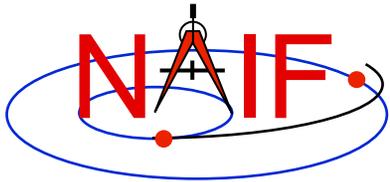
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First, let's go over the important steps in the process of writing a Mice-based program and putting it to work:

- **Understand the geometry problem.**
- **Identify the set of SPICE kernels that contain the data needed to perform the computation.**
- **Formulate an algorithm to compute the quantities of interest using SPICE.**
- **Write and execute the program.**
- **Get actual kernel files and verify that they contain the data needed to support the computation for the time(s) of interest.**
- **Run the program.**

To illustrate these steps, let's write a program that computes the apparent intersection of the boresight ray of a given CASSINI science instrument with the surface of a given Saturnian satellite. The program will compute:

- **Planetocentric and planetodetic (geodetic) latitudes and longitudes of the intercept point.**
- **Range from spacecraft to intercept point.**
- **Illumination angles (phase, solar incidence, and emission) at the intercept point.**



Observation geometry

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We want the boresight intercept on the surface, range from s/c to intercept, and illumination angles at the intercept point.

When? **TIME (UTC, TDB or TT)**

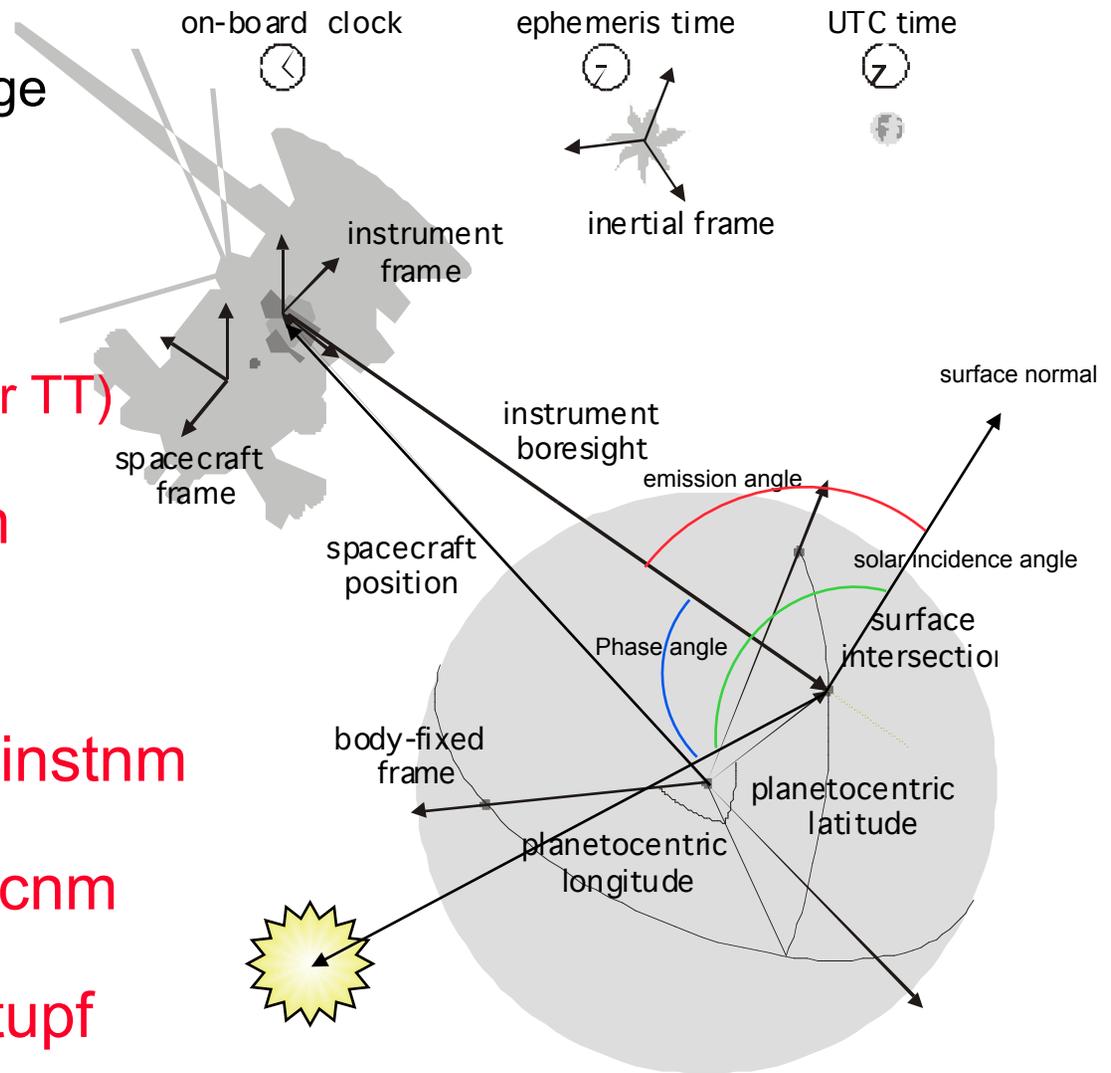
On what object? **satnm**

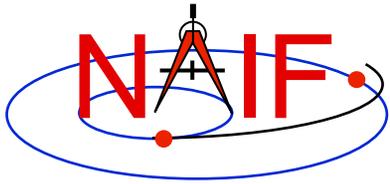
In what frame? **fixref**

For which instrument? **instnm**

For what spacecraft? **scnm**

Using what model? **setupf**





Needed Data

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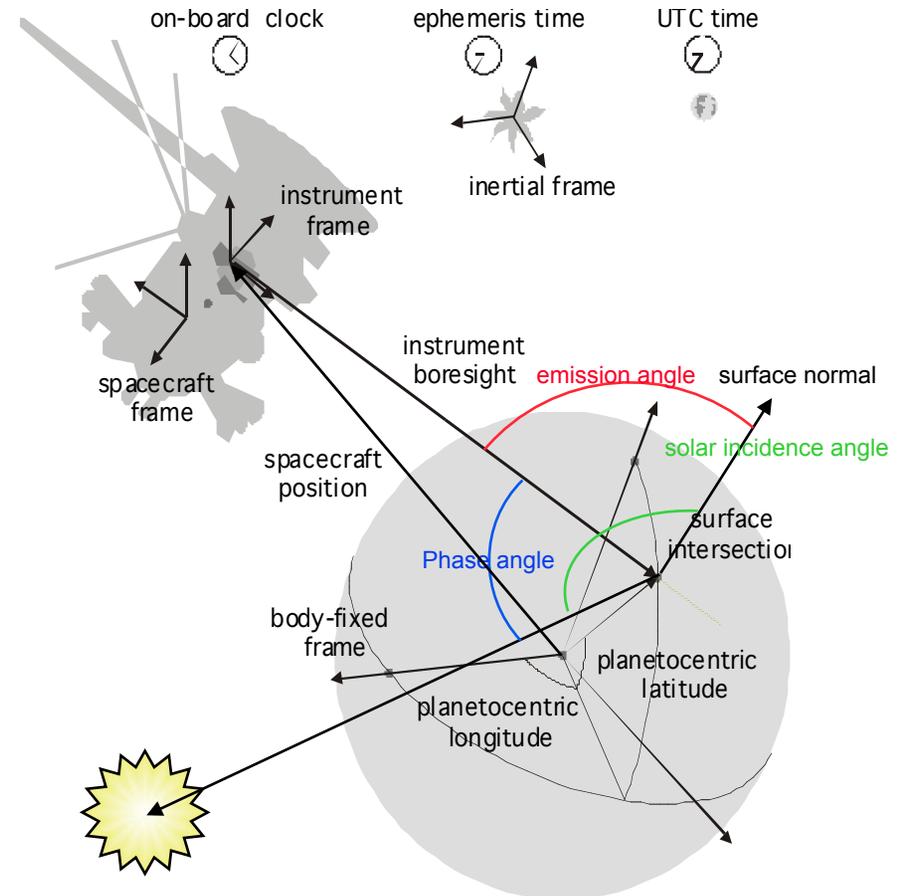
Time transformation kernels

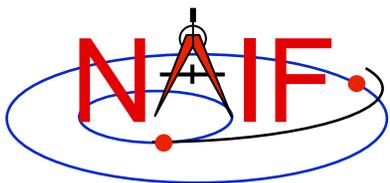
Orientation models

Instrument descriptions

Shapes of satellites, planets

Ephemerides for spacecraft,
Saturn barycenter and satellites.





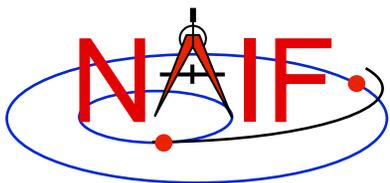
Which Kernels are Needed?

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Data required to compute vectors, rotations and other parameters shown in the picture are stored in the SPICE kernels listed below.

Note: these kernels have been selected to support this presentation; they should not be assumed to be appropriate for user applications.

Parameter	Kernel Type	File name
-----	-----	-----
time conversions	generic LSK	naif0009.tls
	CASSINI SCLK	cas00084.tsc
satellite orientation	CASSINI PCK	cpck05Mar2004.tpc
satellite shape	CASSINI PCK	cpck05Mar2004.tpc
satellite position	planet/sat ephemeris SPK	020514_SE_SAT105.bsp
planet barycenter position	planet SPK	981005_PLTEPH-DE405S.bsp
spacecraft position	spacecraft SPK	030201AP_SK_SM546_T45.bsp
spacecraft orientation	spacecraft CK	04135_04171pc_psiv2.bc
instrument alignment	CASSINI FK	cas_v37.tf
instrument boresight	Instrument IK	cas_iss_v09.ti



Load kernels

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The easiest and most flexible way to make these kernels available to the program is via `cspice_furnsh`. For this example we make a setup file (also called a “metakernel” or “furnsh kernel”) containing a list of kernels to be loaded:

Note: these kernels have been selected to support this presentation; they should not be assumed to be appropriate for user applications.

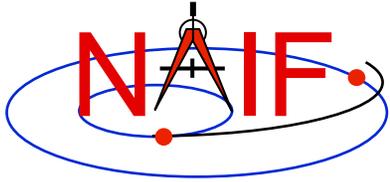
```
\begindata
```

```
KERNELS_TO_LOAD = ('naif0009.tls', 'cas00084.tsc', 'cpck05Mar2004.tpc',  
                   '020514_SE_SAT105.bsp', '981005_PLTEPH-DE405S.bsp',  
                   '030201AP_SK_SM546_T45.bsp', '04135_04171pc_psiv2.bc',  
                   'cas_v37.tf', 'cas_iss_v09.ti')
```

```
\begintext
```

and we make the program prompt for the name of this setup file:

```
setupf = input('Enter setup file name > ', 's');  
cspice_furnsh( setupf )
```

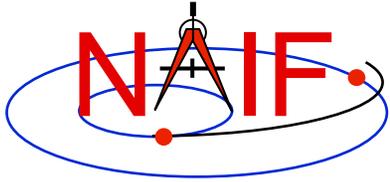


Programming Solution

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- **Prompt for setup file (“metakernel”) name; load kernels specified via setup file. (Done on previous chart.)**
- **Prompt for user inputs required to completely specify problem. Obtain further inputs required by geometry routines via Mice calls.**
- **Compute the intersection of the boresight direction ray with the surface of the satellite, presented as a triaxial ellipsoid.**
- **If there is an intersection:**
 - **Convert Cartesian coordinates of the intersection point to planetocentric latitudinal and planetodetic coordinates**
 - **Compute spacecraft-to-intercept point range**
 - **Find the illumination angles (phase, solar incidence, and emission) at the intercept point**
- **Display the results.**

We discuss the geometric portion of the problem first.



Compute surface intercept

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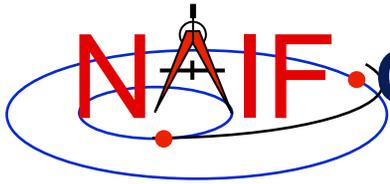
Compute the intercept point (`point`) of the boresight vector (`insite`) specified in the instrument frame (`iframe`) of the instrument mounted on the spacecraft (`scnm`) with the surface of the satellite (`satnm`) at the TDB time of interest (`et`) in the satellite's body-fixed frame (`fixref`). This call also returns the light-time corrected epoch at the intercept point (`trgepc`), the spacecraft-to-intercept point vector (`srfvec`), and a flag indicating whether the intercept was found (`found`). We use "converged Newtonian" light time plus stellar aberration corrections to produce the most accurate surface intercept solution possible. We model the surface of the satellite as an ellipsoid.

```
[point, trgepc, srfvec, found] = cspice_sincpt( ...  
    'Ellipsoid', satnm, et, fixref, 'CN+S', scnm, iframe, insite );
```

The range we want is obtained from the outputs of `cspice_sincpt`. These outputs are defined only if a surface intercept is found. If `found` is true, the spacecraft-to-surface intercept range is the norm of the output argument `srfvec`. Units are km. We use the MATLAB function `norm` to obtain the norm:

```
norm( srfvec )
```

We'll write out the range data along with the other program results.



Compute Lat/Lon and Illumination Angles

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Compute the planetocentric latitude (`pclat`) and longitude (`pclon`), as well as the planetodetic latitude (`pdlat`) and longitude (`pdlon`) of the intersection point.

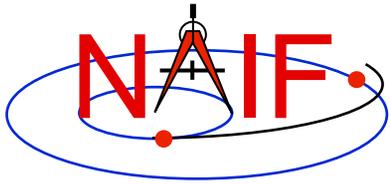
```
if ( found )
    [r, pclon, pclat] = cspice_reclat( point );

    % Let re, rp, and f be the satellite's longer equatorial
    % radius, polar radius, and flattening factor.
    re = radii(1);
    rp = radii(3);
    f = ( re - rp ) / re;

    [pdlat, pdlon, alt] = cspice_recgeo( point, re, f );
```

The illumination angles we want are the outputs of `cspice_ilumin`. Units are radians.

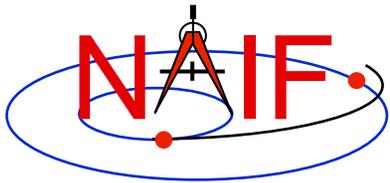
```
[trgepc, srfvec, phase, solar, emissn] = cspice_ilumin( ...
    'Ellipsoid', satnm, et, fixref, 'CN+S', scnm, point );
```



Geometry Calculations: Summary

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```
% Compute the boresight ray intersection with the surface of the
% target body.
[point, trgepc, srfvec, found] = cspice_sincpt( ...
    'Ellipsoid', satnm, et, fixref, 'CN+S', scnm, iframe, insite );
% If an intercept is found, compute planetocentric and planetodetic
% latitude and longitude of the point.
if ( found )
    [r, pclon, pclat] = cspice_reclat( point );
    % Let re, rp, and f be the satellite's longer equatorial
    % radius, polar radius, and flattening factor.
    re = radii(1);
    rp = radii(3);
    f = ( re - rp ) / re;
    [pdlon, pdlat, alt] = cspice_recgeo( point, re, f );
    % Compute illumination angles at the surface point.
    [trgepc, srfvec, phase, solar, emissn] = cspice_ilumin( ...
        'Ellipsoid', satnm, et, fixref, 'CN+S', scnm, point );
    ...
else
    ...
```



Get inputs - 1

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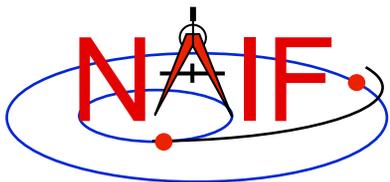
The code above used quite a few inputs that we don't have yet:

- TDB epoch of interest (`et`);
- satellite and s/c names (`satnm`, `scnm`);
- satellite body-fixed frame name (`fixref`);
- satellite ellipsoid radii (`radii`);
- instrument fixed frame name (`iframe`);
- instrument boresight vector in the instrument frame (`insite`);

Some of these values are user inputs; others can be obtained via Mice calls once the required kernels have been loaded.

Let's prompt for the satellite name (`satnm`), satellite frame name (`fixref`), spacecraft name (`scnm`), instrument name (`instnm`) and time of interest (`time`):

```
satnm = input( 'Enter satellite name > ', 's');  
fixref = input( 'Enter satellite frame > ', 's');  
scnm = input( 'Enter spacecraft name > ', 's');  
instnm = input( 'Enter instrument name > ', 's');  
time = input( 'Enter time > ', 's');
```



Get Inputs - 2

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Then we can get the rest of the inputs from Mice calls:

To get the TDB epoch (`et`) from the user-supplied time string (which may refer to the UTC, TDB or TT time systems):

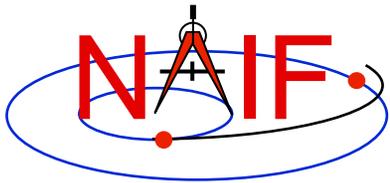
```
et = cspice_str2et( time );
```

To get the satellite's ellipsoid radii (`radii`):

```
radii = cspice_bodvrd( satnm, 'RADII', 3 );
```

To get the instrument boresight direction (`insite`) and the name of the instrument frame (`iframe`) in which it is defined:

```
[instid, found] = cspice_bodn2c( instnm );  
if ( ~found )  
    txt = sprintf( 'Unable to determine ID for instrument: %d', ...  
                  instnm );  
    error(txt)  
end  
  
[shape, iframe, insite, bundry] = cspice_getfov( instid, ROOM );
```



Getting inputs: summary

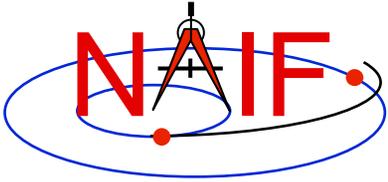
Navigation and Ancillary Information Facility

```
% Prompt for the user-supplied inputs for our program.
setupf = input( 'Enter setup file name > ', 's');
cspice_furnsh( setupf )
satnm   = input( 'Enter satellite name > ', 's');
fixref  = input( 'Enter satellite frame > ', 's');
scnm   = input( 'Enter spacecraft name > ', 's');
instnm = input( 'Enter instrument name > ', 's');
time   = input( 'Enter time > ', 's');

% Get the epoch corresponding to the input time:
et = cspice_str2et( time );

% Get the radii of the satellite.
radii = cspice_bodvrd( satnm, 'RADII', 3 );

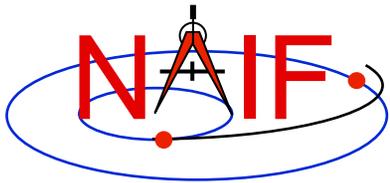
% Get the instrument boresight and frame name.
[instid, found] = cspice_bodn2c( instnm );
if ( ~found )
    txt = sprintf( 'Unable to determine ID for instrument: %d', ...
                  instnm );
    error(txt)
end
[shape, iframe, insite, bundry] = cspice_getfov( instid, ROOM );
```



Display results

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```
...
% Display results. Convert angles from radians to degrees
% for output.
fprintf( 'Intercept planetocentric longitude      (deg):  %11.6f\n', ...
        cspice_dpr()*pclon )
fprintf( 'Intercept planetocentric latitude      (deg):  %11.6f\n', ...
        cspice_dpr()*pclat )
fprintf( 'Intercept planetodetic longitude       (deg):  %11.6f\n', ...
        cspice_dpr()*pdlon )
fprintf( 'Intercept planetodetic latitude       (deg):  %11.6f\n', ...
        cspice_dpr()*pdlat )
fprintf( 'Range from spacecraft to intercept point (km): %11.6f\n', ...
        norm(srfvec) )
fprintf( 'Intercept phase angle                 (deg):  %11.6f\n', ...
        cspice_dpr()*phase )
fprintf( 'Intercept solar incidence angle       (deg):  %11.6f\n', ...
        cspice_dpr()*solar )
fprintf( 'Intercept emission angle             (deg):  %11.6f\n', ...
        cspice_dpr()*emissn )
else
    disp( ['No intercept point found at ' time ] )
end
```



Complete the program

Navigation and Ancillary Information Facility

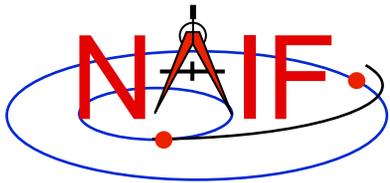
To finish up the program we need to declare the variables we've used.

- We'll highlight techniques used by NAIF programmers
- Add remaining MATLAB code required to make a syntactically valid program

```
ROOM    = 10;
R2D     = cspice_dpr;

% Prompt for the user-supplied inputs for our program.
setupf = input( 'Enter setup file name > ', 's');
cspice_furnsh( setupf )

satnm  = input( 'Enter satellite name > ', 's');
fixref = input( 'Enter satellite frame > ', 's');
scnm   = input( 'Enter spacecraft name > ', 's');
instnm = input( 'Enter instrument name > ', 's');
time   = input( 'Enter time > ', 's');
```



Complete source code - 1

Navigation and Ancillary Information Facility

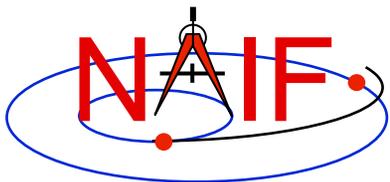
```
% Get the epoch corresponding to the input time:
et = cspice_str2et( time );

% Get the radii of the satellite.
radii = cspice_bodvrd( satnm, 'RADII', 3 );

% Get the instrument boresight and frame name.
[instid, found] = cspice_bodn2c( instnm );

if ( ~found )
    txt = sprintf( 'Unable to determine ID for instrument: %d', ...
                  instnm );
    error(txt)
end

[shape, iframe, insite, bundry] = cspice_getfov( instid, ROOM );
```



Complete source code - 2

Navigation and Ancillary Information Facility

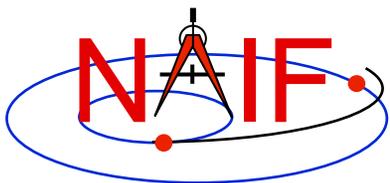
```
% Compute the boresight ray intersection with the surface of the
% target body.
[point, trgepc, srfvec, found] = cspice_sincpt( ...
    'Ellipsoid', satnm, et, fixref, 'CN+S', scnm, iframe, insite );

% If an intercept is found, compute planetocentric and planetodetic
% latitude and longitude of the point.
if ( found )
    [r, pclon, pclat] = cspice_reclat( point );

    % Let re, rp, and f be the satellite's longer equatorial
    % radius, polar radius, and flattening factor.
    re = radii(1);
    rp = radii(3);
    f = ( re - rp ) / re;

    [pdlon, pdlat, alt] = cspice_recgeo( point, re, f );

% Compute illumination angles at the surface point.
[trgepc, srfvec, phase, solar, emissn] = cspice_ilumin( ...
    'Ellipsoid', satnm, et, fixref, 'CN+S', scnm, point );
```



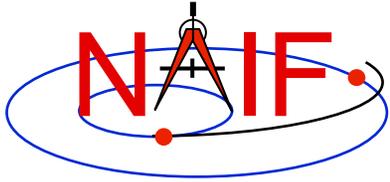
Complete source code - 3

Navigation and Ancillary Information Facility

```
% Display results. Convert angles from radians to degrees
% for output.
fprintf( 'Intercept planetocentric longitude      (deg): %11.6f\n', ...
        R2D*pclon )
fprintf( 'Intercept planetocentric latitude      (deg): %11.6f\n', ...
        R2D*pclat )
fprintf( 'Intercept planetodetic longitude       (deg): %11.6f\n', ...
        R2D*pdlon )
fprintf( 'Intercept planetodetic latitude        (deg): %11.6f\n', ...
        R2D*pdlat )
fprintf( 'Range from spacecraft to intercept point (km): %11.6f\n', ...
        norm(srfvec) )
fprintf( 'Intercept phase angle                  (deg): %11.6f\n', ...
        R2D*phase )
fprintf( 'Intercept solar incidence angle        (deg): %11.6f\n', ...
        R2D*solar )
fprintf( 'Intercept emission angle              (deg): %11.6f\n', ...
        R2D*emissn )

else
    disp( ['No intercept point found at ' time ]
end

% Unload the kernels and clear the kernel pool
cspice_kclear
```



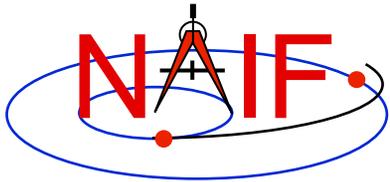
Running the program

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It looks like we have everything taken care of:

- We have all necessary kernels
- We made a setup file (metakernel) pointing to them
- We wrote the program

Let's run it.



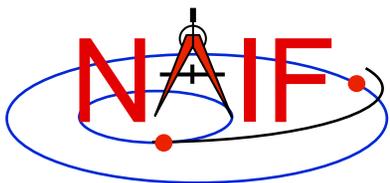
Running the program

Navigation and Ancillary Information Facility

```
Terminal Window

>> prog_geometry
Enter setup file name > setup.ker
Enter satellite name > PHOEBE
Enter satellite frame > IAU_PHOEBE
Enter spacecraft name > CASSINI
Enter instrument name > CASSINI_ISS_NAC
Enter time > 2004 jun 11 19:32:00

Intercept planetocentric longitude (deg): 39.843719
Intercept planetocentric latitude (deg): 4.195878
Intercept planetodetic longitude (deg): 39.843719
Intercept planetodetic latitude (deg): 5.048011
Range from spacecraft to intercept point (km): 2089.169724
Intercept phase angle (deg): 28.139479
Intercept solar incidence angle (deg): 18.247220
Intercept emission angle (deg): 17.858309
```



Backup

Navigation and Ancillary Information Facility

- **Latitude definitions:**

- Planetocentric latitude of a point P: angle between segment from origin to point and x-y plane (red arc in diagram).
- Planetodetic latitude of a point P: angle between x-y plane and extension of ellipsoid normal vector N that connects x-y plane and P (blue arc in diagram).

